Dark matter, neutrino masses and high scale validity of an inert Higgs doublet model

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Abstract

We consider a two-Higgs doublet scenario containing three $SU(2)_L$ singlet heavy neutrinos with Majorana masses. The second scalar doublet as well as the neutrinos are odd under a Z_2 symmetry. This scenario not only generates Majorana masses for the light neutrinos radiatively but also makes the lighter of the neutral Z_2 -odd scalars an eligible dark matter candidate, in addition to triggering leptogenesis at the scale of the heavy neutrino masses. Taking two representative values of this mass scale, we identify the allowed regions of the parameter space of the model, which are consistent with all dark matter constraints. At the same time, the running of quartic couplings in the scalar potential to high scales is studied, thus subjecting the regions consistent with dark matter constraints to further requirements of vacuum stability, perturbativity and unitarity. It is found that part of the parameter space is consistent with all of these requirements all the way up to the Planck scale, and also yields the correct signal strength in the diphoton channel for the scalar observed at the Large Hadron Collider.

1 Introduction

The discovery of 'a Higgs-like boson' [1, 2], apparently completes the Standard electroweak model (SM). However, it also compels us to mull over the necessity of physics beyond the SM. To mention one motivation for this, a SM Higgs of mass around 125 GeV, can break the absolute stability of vacuum at 10^{8-9} GeV, if the top quark mass (M_t) and the strong coupling constant (α_s) are on the upper sides of their respective uncertainty bands. A recent next-to-next-to-leading order (NNLO) study [3] reveals that absolute stability up to the Planck scale requires

$$M_h[{\rm GeV}] > 129.4 + 1.4(\frac{M_t[{\rm GeV}] - 173.1}{0.7}) - 0.5(\frac{\alpha_s({\rm M_Z}) - 0.1184}{0.0007}) \pm 1.0_{\rm th}$$
 (1)

which, with all theoretical and experimental errors in M_t and α_s , implies that the absolute vacuum stability of SM is excluded at 95% CL for $M_h < 126$ GeV. To compensate the negative contribution

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from the fermion loops (mainly the top quark loop), which rapidly drives the quartic self-coupling of Higgs to below zero, we need extra scalar loops that can ameliorate the problem. This provides one of the important motivations to consider a theory with new scalars which can make the vacuum stable up to high scales while having consistent results at low scales.

At the same time it is worthwhile to remember two rather pressing issues which prompt one to look beyond the Standard Model (BSM). These are the non-zero mass and mixing of neutrinos and the likely existence of Weakly Interacting Massive Particle(WIMP), contributing to Dark Matter (DM). While various Solar, atmospheric, reactor and accelerator-based experiments provide strong evidence for non-zero masses of neutrinos and mixing among their different flavors, several astrophysical and cosmological observations suggest the existence of some exotic particles that constitute the DM content of the universe. Given all this, it is an attractive idea to look for a theory that can simultaneously address all of the aforesaid problems in one framework. Here we consider one such scenario.

We investigate a model, first proposed in reference [4], that extends the SM with an extra Higgs doublet and three right handed neutrinos with a Z_2 symmetry, under which all SM particles are even while this additional scalar doublet and the right-handed neutrinos are odd. This symmetry prevents the additional doublet from having a vacuum expectation value(vev) thus vetoing the tree-level neutrino mass generation. Moreover, there exists a stable scalar particle in the form of the lightest neutral mass eigenstate of the additional doublet, which yields an eligible DM candidate. The extra doublet is essentially an Inert Doublet. Although a lot of study has already taken place on minimal Inert Doublet models [5–23], the extra appeal of this model lies in the radiative generation of neutrino mass. Though various aspects of this scenario have already been investigated [24–35], the present study includes the following points which have not been emphasized before.

- We investigate the vacuum stability of this model at various scales, and identify the regions of its parameter space, which keeps the model valid all the way unto the Planck scale. The contribution of additional scalar fields as well as the right-handed neutrinos to the renormalisation group (RG) equations [given in Appendix A] has been taken into account here. Using these modified RG equations we evaluate the scalar quartic couplings at different scales. During the evolution of the quartic couplings we demand not only vacuum stability but also perturbativity of the couplings as well as unitarity of the 2 → 2 scattering matrix at each scale.
- The heavy right-handed neutrinos introduce a new mass scale (M) to the theory and the neutrino Yukawa couplings contribute to the RG evolution only at this mass scale. This brings out greater implications on the parameter space that distinguishes this model from minimal Inert doublet models. In this paper, we will show the salient features of the model that emerges from the above fact.
- This model also contributes to the study of leptogenesis due to this heavy right-handed fermions. The values of the right-handed neutrino mass scale used in the high-scale analysis are taken to be commensurate with leptogenesis constraints. Thus the part of the model space consistent unto the Planck scale is also supportive of leptogenesis.
- We also examine the candidature of the lightest Z_2 -odd particle as DM candidate, and identify the allowed values of the couplings yield the right relic abundance. We ensure that the DM candidate is consistent with the recent result of direct detection experiments. We identify a substantial region of the parameter space, which simultaneously satisfy the vacuum stability,

perturbativity and unitarity requirements and accommodate the DM candidate with the correct relic density.

• Finally, we examine the 125 GeV scalar and make sure that the signal strengths in the observed channels (such as diphotons) are consistent with data from the Large Hadron Collider (LHC).

We organize the paper as follows. In Section 2, we briefly describe the model and its various features. In Section 3, we explain all the theoretical constraints and collider constraints that we use in the RG running of different quartic couplings. Next, in Section 4, we discuss the DM aspects of this model. After explaining our analysis strategy in Section 5, we present our results related to high-scale validity in Section 6. Finally in Section 7, we summarise our results.

2 The Radiative Neutrino Mass Model with an Inert Doublet

In addition to the SM fields, the radiative neutrino mass model with an inert doublet [4], contains a Higgs doublet (Φ_2) and three right handed (RH) neutrinos (N_i) with an unbroken Z_2 symmetry, under which the doublet and the right handed neutrinos are odd while all other SM particles are even. Being odd under the symmetry, Φ_2 does not acquire any vacuum expectation value (vev) and has no tree-level couplings to fermions.

The relevant Yukawa and mass terms are

$$-\mathcal{L}_{Y} = y_{ij}\bar{N}_{i}\tilde{\Phi}_{2}^{\dagger}\ell_{j} + h.c + \frac{M_{i}}{2} \left(\bar{N}_{i}^{c}N_{i} + h.c\right), (i, j = 1 - 3)$$
(2)

Here ℓ_i are the left-handed lepton doublets and M_i are the Majorana mass term for the heavy right-handed neutrinos N_i

The scalar potential is

$$V = \lambda_{1}(\Phi_{1}^{\dagger}\Phi_{1})^{2} + \lambda_{2}(\Phi_{2}^{\dagger}\Phi_{2})^{2} + \lambda_{3}(\Phi_{1}^{\dagger}\Phi_{1})(\Phi_{2}^{\dagger}\Phi_{2}) + \lambda_{4}(\Phi_{2}^{\dagger}\Phi_{1})(\Phi_{1}^{\dagger}\Phi_{2})$$

$$+ \left[\frac{\lambda_{5}}{2}(\Phi_{1}^{\dagger}\Phi_{2})^{2} + \text{h.c.}\right] + m_{\Phi_{1}}^{2}\Phi_{1}^{\dagger}\Phi_{1} + m_{\Phi_{2}}^{2}\Phi_{2}^{\dagger}\Phi_{2}$$

$$(3)$$

where all parameters are real, and Φ_1 is the SM Higgs doublet.

The two scalar doublets can be written as

$$\Phi_1 = \begin{pmatrix} G^+ \\ \frac{1}{\sqrt{2}}(v+h+iG) \end{pmatrix} \text{ and, } \Phi_2 = \begin{pmatrix} H^+ \\ \frac{1}{\sqrt{2}}(H+iA) \end{pmatrix}$$
 (4)

where, v = 246 GeV, is the electroweak vacuum expectation value (vev). One thus has five physical states (h, H, A, H^{\pm}) and three Goldstone bosons (G, G^{\pm}) . While h corresponds to the physical SM-like Higgs field, the inert doublet components are one CP-even neutral scalar(H), one CP-odd neutral scalar(A) and a pair of charged scalars (H^{\pm}) . The physical masses are given by

$$M_{H^{\pm}}^{2} = m_{\Phi_{2}}^{2} + \frac{1}{2}\lambda_{3}v^{2}$$
$$M_{H}^{2} = m_{\Phi_{2}}^{2} + \lambda_{L}v^{2}$$

$$M_A^2 = m_{\Phi_2}^2 + \lambda_A v^2 \tag{5}$$

where $\lambda_{L/A} = \frac{1}{2}(\lambda_3 + \lambda_4 \pm \lambda_5)$. The value of λ_1 is determined using $M_h = 125$ GeV.

Majorana masses for the light neutrinos are generated radiatively through one-loop exchange of the \mathbb{Z}_2 -odd neutral scalars. The general expression for the loop-induced contributions to the light neutrino mass matrix [4] is

$$\mathcal{M}_{ij}^{\nu} = \sum_{k=1}^{3} \frac{y_{ik} y_{jk} M_k}{16\pi^2} \left[\frac{M_H^2}{M_H^2 - M_k^2} \ln \frac{M_H^2}{M_k^2} - \frac{M_A^2}{M_A^2 - M_k^2} \ln \frac{M_A^2}{M_k^2} \right]$$
(6)

Thus, the neutrino masses and mixing are determined by the inert scalar masses and the right-handed neutrino masses M_i . These masses represent the scale of lepton number violation and hence that of leptogenesis [36,37] in this model. Some studies have already been done in this context [11,30,32,33,38]. Our choice of the right-handed neutrino (Majorana) mass scales made our study automatically compatible with leptogenesis. To satisfy the necessary constraints in the low DM mass region where $M_{DM} < M_W$, one must take the lightest of the Majorana masses to be $M \ge 110$ TeV, where as in the high DM mass region ($M_{DM} > 500$ GeV) the bound is only 1 TeV [11]. Hence, to be consistent in both cases, we use two values of M, (a) M = 110 TeV and (b) $M = 10^9$ TeV. While, in one hand, choice of (a) is motivated by the idea of having the lowest possible leptogenesis scale, we choose to work with (b) which have interesting consequences on the RG runnning. We will show in later sections, how the mass scale (b) of M affects the stability of the vacuum mainly in the high DM mass region and eventually explain the physical reasons behind it.

Along with the above restrictions we also demand $M_{\nu} \sim \mathcal{O}(0.1 \text{ eV})$ to be consistent with neutrino oscillation data for some fixed M and other exotic scalar masses. However, for simplicity, we consider only one diagonal Yukawa coupling (y_{ν}) and do not look into the hierarchical details of the Yukawa matrix. At this point, it is to be noted that a more rigorous study with the intricate flavor structure of the neutrino Yukawa matrix can highlight the region of the model space that fits the observed pattern of neutrino mixing [35]. However, we would like to emphasize that the broad conclusions on the high-scale validity of this scenario vis-a-vis the DM constraints remain unchanged. Finally, we should mention that the lighter state between H and A is the DM candidate. We present our illustrative results for cases where H plays this role.

3 Constraints from perturbativity, unitarity, vacuum stability and collider data

In this section, we briefly describe the constraints that are imposed on the model parameters and how exactly they shape the results so obtained.

3.1 Vacuum stability

The scalar potential is considered bounded from below, if it does not turn negative for large field values along all possible field directions. In this model, stability of the electroweak vacuum is ensured up to some specified energy scale if the following conditions are satisfied for all scales Q up to that scale:

$$vsc1 : \lambda_1(Q) > 0 \tag{7a}$$

$$vsc2 : \lambda_2(Q) > 0 \tag{7b}$$

vsc3 :
$$\lambda_3(Q) + \sqrt{\lambda_1(Q)\lambda_2(Q)} > 0$$
 (7c)

$$vsc4 : \lambda_3(Q) + \lambda_4(Q) - |\lambda_5(Q)| + \sqrt{\lambda_1(Q)\lambda_2(Q)} > 0$$
 (7d)

Such conditions have been elaborately discussed in literature [39–42]. One should make a note that these conditions ensure *absolute* stability of the electroweak vacuum. For metastability, the conditions are somewhat less stringent.

3.2 Perturbativity

For the scalar quartic coupling $\lambda_i (i = 1 - 5)$, the criterion for perturbativity is,

$$\lambda_i(Q) < 4\pi \tag{8}$$

The corresponding constraints for the Yukawa and gauge interactions are,

$$y_i(Q), \ g_i(Q) < \sqrt{4\pi} \tag{9}$$

where, Q represents the energy scale at which they are being computed. We demand that the criteria in Eq. 8 be obeyed at all energy scales up to the cut-off of this model.

3.3 Unitarity

A further set of conditions come on demanding unitarity of the scattering matrix comprising all $2 \to 2$ channels involving, by the optical theorem [43–46]. In our context, this translates into the condition that each distinct eigenvalue of the aforementioned amplitude matrix be bounded above at 8π (after factoring out $\frac{1}{16\pi}$ from the matrix). These eigenvalues are:

$$a_{\pm} = \frac{3}{2}(\lambda_{1} + \lambda_{2}) \pm \sqrt{\frac{9}{2}(\lambda_{1} - \lambda_{2})^{2} + (2\lambda_{3} + \lambda_{4})^{2}}$$

$$b_{\pm} = \frac{1}{2}(\lambda_{1} + \lambda_{2}) \pm \sqrt{\frac{1}{4}(\lambda_{1} - \lambda_{2})^{2} + \lambda_{5}^{2}},$$

$$c_{\pm} = d_{\pm} = \frac{1}{2}(\lambda_{1} + \lambda_{2}) \pm \sqrt{\frac{1}{4}(\lambda_{1} - \lambda_{2})^{2} + \lambda_{5}^{2}},$$

$$e_{1} = (\lambda_{3} + 2\lambda_{4} - 3\lambda_{5}),$$

$$e_{2} = (\lambda_{3} - \lambda_{5}),$$

$$f_{1} = f_{2} = (\lambda_{3} + \lambda_{4}),$$

$$f_{+} = (\lambda_{3} + 2\lambda_{4} + 3\lambda_{5}),$$

$$f_{-} = (\lambda_{3} + \lambda_{5}).$$
(10)

3.4 Collider data

In addition to the theoretical constraints discussed above, important bounds on scalar mass parameters come from collider data.

- In order to identify h with the SM-like Higgs as observed by the ATLAS and CMS collaborations, one requires $M_h \simeq 125$ GeV.
- To be consistent with the LEP bounds, one must have

$$M_H + M_A > M_Z$$
, (11)
 $M_{H^{\pm}} + M_{A/H} > M_W$.

Moreover, for neutralino as in the supersymmetric context, LEP-II searches limit pseudo-scalar mass (M_A) to 100 GeV when $M_H < M_A$ [8, 47]. Similarly, chargino search data, properly extrapolated, imply $M_{H^{\pm}} > 70 \text{ GeV } [48].$

• Though all the tree-level couplings of h are identical to those of the SM Higgs, the charged scalar H^{\pm} potentially modifies the loop induced couplings $h\gamma\gamma$ and $hZ\gamma$ via loop contributions [49–53]. We theoretically compute the signal strength $\mu_{\gamma\gamma}$ for h decaying to the diphoton channel as the ratio of the decay width in the 'model' to that in the SM. Its experimental value reported by the ATLAS and CMS collaborations stand at 1.17 ± 0.27 and 1.13 ± 0.24 respectively [54, 55]. Demanding the signal strength to be within 2σ limits of the experimentally quoted values puts further constraints on the model. We use the limit on $\mu_{\gamma\gamma}$ weighted as below:

$$\frac{1}{\sigma^2} = (\frac{1}{\sigma^2})_{ATLAS} + (\frac{1}{\sigma^2})_{CMS}$$

$$\frac{\mu_{\gamma\gamma}}{\sigma^2} = (\frac{\mu_{\gamma\gamma}}{\sigma^2})_{ATLAS} + (\frac{\mu_{\gamma\gamma}}{\sigma^2})_{CMS}$$
(12)

$$\frac{\mu_{\gamma\gamma}}{\sigma^2} = (\frac{\mu_{\gamma\gamma}}{\sigma^2})_{ATLAS} + (\frac{\mu_{\gamma\gamma}}{\sigma^2})_{CMS} \tag{13}$$

where, the numerators on the right hand side denotes the central values of the respective experimental results and σ are the corresponding uncertainties.

Dark Matter Issues 4

As stated earlier, we identify H as the DM candidate. For complimentary choice, namely, A with the same mass as the DM candidate, we have checked that the contribution to the relic density is of similar magnitude. The relevant DM constraints to be considered are as follows:

• According to recent PLANCK experiment [56] the observed cold DM relic density is

$$\Omega_{\rm DM}h^2 = 0.1199 \pm 0.0027\tag{14}$$

We restrict our result up to 3σ deviation from the central value.

- Strong constraints come from direct DM searches. Recently, XENON100 [57] and LUX [58] experiments have put upper-bound on the DM-nucleon scattering cross-section for a wide range of the DM mass. In our case, the direct detection cross-section strategy is based on t-channel Higgs mediation. We choose to work in the region of the parameter space allowed by the LUX limit.
- For $M_H \leq M_h/2$, the decay mode of Higgs to two DM particle $(h \to HH)$ will presumably contribute to the invisible decay width of the Higgs boson. We take into account the current constraint on the Higgs invisible branching ratio (< 20%) from model independent Higgs precision analysis [59, 60].

• Our parameter scan has included situations where the Z_2 -odd scalar and the pseudo-scalar are close by in mass and with co-annihilation yield correct relic abundance within the $\pm 3\sigma$ of the observed values when the DM mass is high (> 500 GeV). For DM mass in between 100 GeV and 500 GeV, there is prima facie the possibility of co-annihilation between the spinless DM candidate and a right-handed neutrino which can give rise to correct relic abundance, as discussed in Ref. [61]. In our case, however, such a possibility is ruled out by the additional requirement of leptogenesis, which in turn puts a lower bound on the mass range for the right-handed neutrino(s).

Hence, only the following two cases survive all the above constraints.

4.1 Case-A: Low mass DM ($50 \text{ GeV} < M_H < 90 \text{ GeV}$)

In this mass region, the dominant annihilation channel for H goes to the DM self-annihilation processes through h mediation. This keeps the relic density at the right level. For both positive and negative values of DM-DM-Higgs couplings, the relic density remains within the allowed range as long as the $M_H < 90$ GeV. The sub-dominant contribution to the relic density comes from the t-channel processes to vector boson final states mediated by A and H^{\pm} . A detailed discussion in this regard on a similar model can be found, for example in [11,13,22]. However, the coupling λ_2 has no effect in the relic density calculation. In the next section, we will discuss this results elaborately. One more notable point is that, annihilation processes that mediated by the heavy right-handed neutrinos give negligible contribution (less than 1%) to the relic density calculation. These processes are suppressed by the heavy mediator mass.

4.2 Case-B: High mass DM $(M_H > 500 \text{ GeV})$

The interesting feature of this region is that, the correct relic abundance can be achieved if and only if H, A and H^{\pm} are almost degenerate, at most have a mass difference of the order of 10 GeV. This is mainly because at this high mass, the annihilation channels with vector boson final states can have direct quartic couplings $(HHVV, V = W^{\pm}, Z)$ or can be mediated by any of the three scalars through t/u channels. These diagrams yield too large an annihilation cross-section to match with the proper relic density. However, cancellation between direct quartic coupling diagrams and t/u channels diagram occurs for mass-degenerate H, A and H^{\pm} , which in turn brings down the annihilation cross-section to the desired range.

5 Analysis strategy

The aim of this study is to probe the parameter space of an inert doublet model (IDM) augmented with heavy RH neutrinos compatible with various theoretical and experimental constraints elaborated in the previous sections. We carry out our investigation in the two separate mass regions. In each region, we scan over the relevant parameters, namely, the masses M_H , M_A and M_H^{\pm} , and the coupling λ_L . With M_h fixed at 125 GeV, λ_2 , λ_L , M_H , M_A and M_H^{\pm} fix all the remaining quartic interactions. At this point it is to be mentioned that the parameter λ_2 can not be constrained by any physical observable at least at tree level. However, λ_2 is bounded by the stability and perturbativity condition

and we have explicitly checked that $0 < \lambda_2 < 0.36$ to satisfy the theoretical constraints. Rather than scanning over λ_2 , we have demonstrated our results with two benchmark values (0.1 and 0.001) for it, both of which are well within the above limit.

These quartic couplings are then used as the electroweak boundary conditions at $Q = M_t$ and their RG evolution to high scales is studied. Here M_t denotes the top quark pole mass. The reader is reminded that the effect of the RH neutrinos is turned on at a scale Q = M. Thus, for $M_t \leq Q \leq M$, we do not include the RH neutrino contributions to the one-loop RG equations. We include such conditions for $Q \geq M$ and use the RG equations listed in the appendix. The scale up to which the scenario remains consistent is denoted by Λ_{UV} . For a generic parameter point $\lambda_i(Q = M_t)$, we check the aforementioned theoretical constraints at all intermediate scales up to Λ_{UV} . If the constraints are all satisfied, we tag $\lambda_i(Q = M_t)$ as an allowed point. This marks out an allowed region in the parameter space defined at the electroweak scale. Moreover, the effects of constraints stemming from the DM observables and collider searches are examined independently in this region. The finally allowed parameter regions are thus identified and presented for benchmark values of λ_2 and M. We use the publicly available package micrOMEGAs-v3.6.9.2 [62] for DM analysis. In all the cases, we also make sure that our allowed parameter space does not violate the oblique T-parameter constraints [63,64].

Amongst the SM fermions, only the top quark plays the dominant role in the evolution of the couplings. The boundary condition for its Yukawa interaction at the electroweak scale is fixed by $y_t(M_t) = \frac{\sqrt{2}M_t}{v}(1 - \frac{4}{3\pi}\alpha_s(M_t))$. We have used $M_t = 173.39$ GeV throughout our analysis.

6 Numerical results

6.1 Low mass DM region

We perform a detailed parameter space scan where $M_H < 100$ GeV. In this scan, we impose the LEP bounds as discussed in Sec 3.4.

$$\lambda_L : [-1.0, 1.0] \tag{15}$$

$$M_H: [50.0 \text{ GeV}, 90. \text{ GeV}]$$
 (16)

$$M_A: [100. \text{ GeV}, 500. \text{ GeV}]$$
 (17)

$$M_{H^+}: [100. \text{ GeV}, 500. \text{ GeV}]$$
 (18)

We solve the RG equations for two values of M, 110 TeV and 10⁹ TeV respectively. We then show the allowed parameter space in the $\lambda_L - M_H$ plane for different choices of Λ_{UV} in Fig 1 and 2. The regions denoted by A (red), B (cyan) and C (green) denote $\Lambda_{UV} = 10^6$, 10^{16} and 10^{19} GeV respectively. We overlay the region allowed by the Higgs to diphoton signal strength within 2σ limits of the current data as grey region named D. As mentioned earlier, the full analysis is done for two values of λ_2 at the electroweak scale (0.1 and 0.001).

Let us briefly summarize the features that emerges from the figures and their possible explanations.

• From Fig.1, it can be seen that an Inert doublet model with right handed neutrinos posses stable vacuum even up to the Planck scale, but the higher the cut-off scale Λ_{UV} , the less amount of

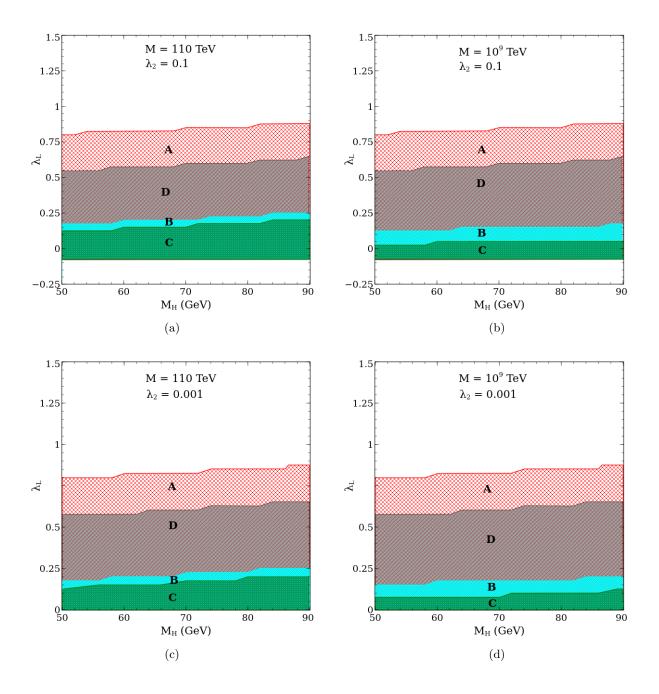


Figure 1: Regions compatible with the theoretical constraints in the M_H - λ_L plane for M=110 TeV (left panel) and 10^9 TeV (right panel) with three different choices of Λ_{UV} and two values of λ_2 (upper and lower panel). The regions denoted by A (red), B (cyan) and C (green) obey these constraints up to $\Lambda_{UV}=10^6$, 10^{16} and 10^{19} GeV, respectively. The grey region denoted by D keeps the Higgs to diphoton signal strength within 2σ limits of the current data.

parameter space becomes allowed. To understand this, one can recollect that the top quark Yukawa coupling in the SM is responsible for the downward evolution of the scalar self-coupling, which poses a threat to the vacuum stability. The presence of the additional scalar quartic couplings ($\lambda's$) in a model like this offsets such an effect; however, the boost thus provided to these couplings tend to violate the perturbativity and unitarity condition. This necessitates a tightrope walking, and the scale up to which it is possible is Λ_{UV} . Hence, it is natural that for higher Λ_{UV} , fewer combinations of parameters will achieve this fine balance, and consequently the allowed region shrinks.

- Fig.1 also shows that λ_L is bounded on both sides and the limits stay almost same for different right handed neutrino mass scale(M). This is of no surprise and can be easily understood. Since $\lambda_3 = \frac{2}{v^2}(M_{H^\pm}^2 M_H^2 + \lambda_L v^2)$, the upper bound on λ_L is imposed by the requirement of perturbative unitarity. This is because a higher positive value of λ_L makes λ_3 large at the electroweak scale which violates perturbative unitarity in the course of its evolution under RG. On the other hand, a large negative value of λ_L induces a large negative value to λ_3 . As a consequence, the vacuum stability condition vsc4 of Eq.7d is violated even near the electroweak(EW) scale. This puts a lower limit on λ_L independent of M and Λ_{UV} , as evident from Fig.1. However, it must be noted that the lower limit of λ_L is not independent of λ_2 , which is another consequence of Eq.7d that requires the condition $\lambda_3 + \lambda_4 |\lambda_5| > -\sqrt{\lambda_1 \lambda_2}$ to be satisfied. With the decrease in the value of λ_2 the lower limit of the combination $\lambda_3 + \lambda_4 |\lambda_5|$ enhances. This combination can be approximated to λ_L in the perturbative regime and hence lower values of λ_2 increases the lower bound of λ_L , as can be seen from the upper and lower panel of Fig.1 that corresponds to $\lambda_2 = 0.1$ and $\lambda_2 = 0.001$ respectively.
- To check the compatibility of the DM constraints with the theoretical ones, we look for the region allowed by the 3σ limits on $\Omega_{\rm DM}h^2$ from PLANCK data and 90% CL limit on the spin-independent DM-nucleon scattering cross-section from LUX data. In Fig. 2, we show the parameter space allowed by the entire set of DM constraints. An inspection of Fig. 1 and Fig. 2 shows that almost the full parameter space allowed by the DM constraints lies in the region which is also favoured by the vacuum stability condition all the way up to the Planck scale. However, the region corresponding to $\lambda_L \leq -0.1$ (shaded region in Fig. 2) does not lead to the stable vacuum.
- It should be noted that in Fig. 2 we keep M_H^{\pm} and M_A fixed at 200 GeV. For higher values of $M_{H^{\pm}}$ and M_A , there is hardly any change in the annihilation cross-section. However, for values of M_H^{\pm} and M_A less than 200 GeV, the allowed region of Fig. 2 gets slightly modified. For example, for $M_H \simeq 70$ GeV and $M_{H^{\pm}} = M_A = 200$ GeV, the DM-DM-Higgs coupling $\lambda_L \simeq 0.007$, but for $M_{H^{\pm}} = M_A = 100$ GeV, one needs $\lambda_L \simeq 0.009$ to satisfy the relic density constraint. It should be noted however that both of the above points in the parameter space are within the stability region as shown in Fig. 1.
 - Therefore, it is not possible to constrain M_H^{\pm} and M_A using DM constraints alone, theoretical constraints however predict strong upper bounds on these masses, as is evident from Fig. 3 and Fig. 4.
- In Fig. 3 and Fig. 4, we show the theoretically valid regions in the $\lambda_3 M_H^{\pm}$ and $\lambda_L M_A$ planes. As in previous cases, we exhibit the same for two different M and λ_2 values. Also, in each case, we overlay the parameter spaces allowed by the 2σ limit of Higgs to diphoton signal strength (region **D**). We observe that a tight upper bound of 160-170 GeV is realised on the masses M_H^{\pm} and M_A for the cut-off at the Planck scale and the couplings (λ_3, λ_L) are also bounded. The upper bounds on the masses M_H^{\pm} and M_A are imposed by the requirement of perturbativity and unitarity up to the desired cut-off. With M_H in the aforementioned range, large masses of the

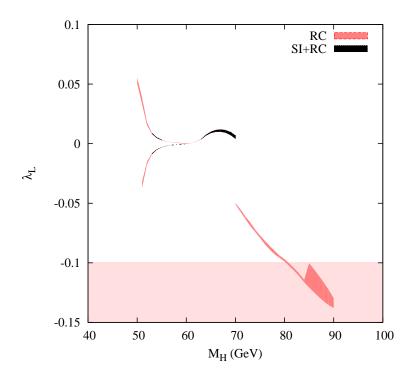


Figure 2: Region allowed by imposing the constraints on relic density(RC) and spin-independent cross-section(SI) for DM-nucleon scattering. The red(gray) region is allowed only by the requirement of $\Omega_{\rm DM}h^2$ being in the correct range. The black region is allowed by both the $\Omega_{\rm DM}h^2$ and direct detection constraints. The shaded horizontal band below is disallowed by vacuum stability conditions. Here, $M_{H^\pm} = M_A = 200$ GeV.

other Z_2 -odd scalars imply high values of the quartic couplings at the electroweak scale which potentially violate perturbativity or unitarity at some high scale. On the other hand, the vacuum stability condition $\lambda_3 + \sqrt{\lambda_1 \lambda_2} \geq 0$ forbids large negative values of λ_3 and perturbativity puts the upper bound. The explanations of the limits of λ_L have already been mentioned above.

• The new physics contribution to the Higgs to diphoton decay channel comes from the charged Higgs loop which is a function of $M_{H^{\pm}}$ and λ_3 . The fact that λ_3 can not be large negative results in a fall in the signal strength for diphoton channel. However, the parameter space for $\Lambda_{UV} = 10^{19}$ GeV is still allowed by the current limits on $\mu_{\gamma\gamma}$ defined in Eq. 13.

6.2 High mass DM region

This section demonstrates the high scale validity of our scenario in the limit of a high DM mass. As discussed earlier, one needs to tune the mass splitting amongst H, H^{\pm} and A and the coupling λ_L to an appropriate degree in order to achieve a relic density within the desired bounds. It is seen that the maximal allowed splitting (ΔM) amongst the masses of the Z_2 odd scalars is 10 GeV. As previously mentioned, for each chosen values of λ_2 (0.1 and 0.001), one is thus motivated to scan the high DM

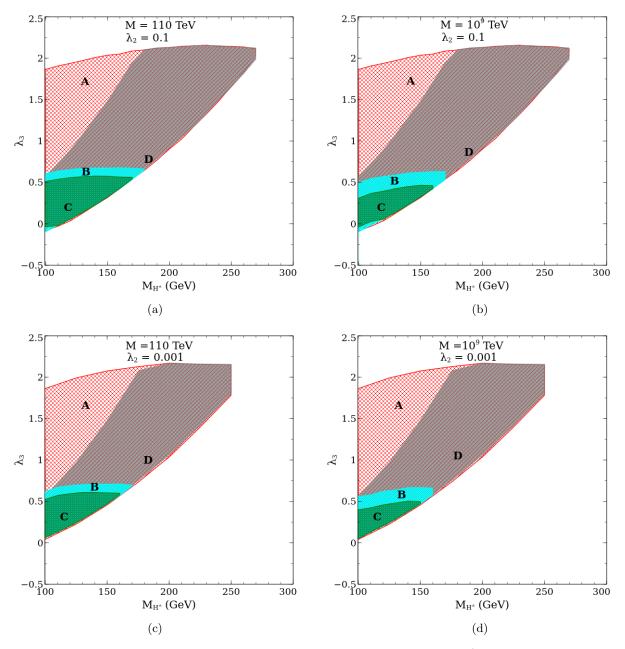


Figure 3: Regions allowed by the theoretical constraints projected in the M_H^{\pm} - λ_3 plane for M=110 TeV (left panel) and 10^9 TeV (right panel) with three different choices of Λ_{UV} and two values of λ_2 (upper and lower panel). The regions denoted by A (red), B (cyan) and C (green) obey those constraints up to $\Lambda_{UV}=10^6$, 10^{16} and 10^{19} GeV, respectively. The grey region denoted by D shows the 2σ allowed limit of the Higgs to diphoton signal strength.

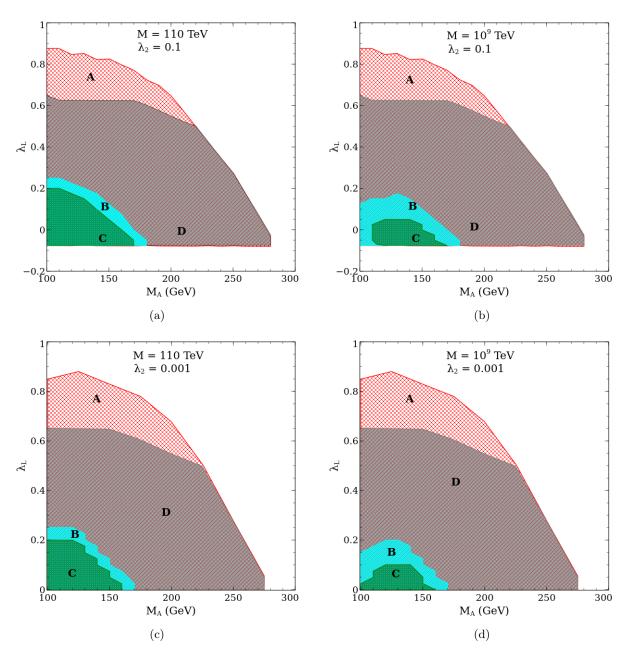


Figure 4: Regions allowed by the theoretical constraints projected in the M_A - λ_L plane for M=110 TeV (left panel) and 10^9 TeV (right panel) with three different choices of Λ_{UV} and two values of λ_2 (upper and lower panel). The regions denoted by A (red), B (cyan) and C (green) obey those constraints up to $\Lambda_{UV}=10^6$, 10^{16} and 10^{19} GeV, respectively. The grey region denoted by D shows the 2σ allowed limit of the Higgs to diphoton signal strength.

mass region in the following ranges:

$$\lambda_L : [-1.0, 1.0] \tag{19}$$

$$M_H: [550.0 \text{ GeV}, 1000.0 \text{ GeV}]$$
 (20)

$$M_A: [M_H, M_H + 10.0 \text{ GeV}]$$
 (21)

$$M_{H^+}: [M_H, M_H + 10.0 \text{ GeV}]$$
 (22)

Unlike the previous case, while the theoretical constraints ruled out a large portion of the parameter space, the DM constraints put a less stringent bound on it in this high DM mass region. Therefore, in Fig. 5 and Fig. 6, we demonstrate the parameter region allowed only by the DM constraints and then overlay the part which are further allowed by theoretical constraints. The full (pink) region denoted by $\mathbf{RC} + \mathbf{SI}$ shows the valid parameter space allowed by DM constraints. In accordance with previous notation, the regions denoted by A (red), B (cyan) and C (green) denote $\Lambda_{UV} = 10^6$, 10^{16} and 10^{19} GeV respectively.

Let us explain the various features of the model that emerges from the figures, in detail.

- (1) The interplay of the theoretical and experimental constraints is studied in the form of correlation plots in the M_H λ_L (Fig. 5) as well as M_{H^\pm} λ_3 (Fig. 6) plane. As can be seen, a significant amount of parameter space is forbidden for the theory to be valid until the Planck scale.
- (2) Similar to the previous section, the upper and lower bounds on λ_L are placed from the requirements of perturbative unitarity and vacuum stability conditions respectively, as shown in Fig. 5. In this region too the lower limit of λ_L has dependence on λ_2 . The same happens for λ_3 , as depicted in Fig. 6.
- (3) It is to be noted that the theoretically allowed parameter space in this high DM mass region is fully consistent with the Higgs to diphoton LHC data. Since, a heavy H^{\pm} which naturally arises in this region, does not cause any significant deviation in diphoton signal strength for the SM-like scalar. This occurs even with a large positive λ_3 (within the bounds shown in Fig. 6). In principle, a large negative λ_3 could modify $\mu_{\gamma\gamma}$ unacceptably. However, as Fig. 6 shows, such values are inconsistent with the aforementioned theoretical constraints.
- (4) It is worth noting here that the parameter space valid until the Planck scale and corresponding to M=110 TeV shrinks significantly for $M=10^9$ TeV, as can be seen in Fig. 5 and Fig. 6. This is because, the significant part of the parameter space is discarded by the stability condition at large M where the neutrino Yukawa couplings becomes large $\mathcal{O}(10^{-1})$. Such large Yukawa coupling contributes to the beta function of λ_2 as shown in Eq. 24b through the terms $+\lambda_2 y_{\nu}^2$ and $-y_{\nu}^4$ that either makes λ_2 perturbative in some cases or λ_2 negative and the vacuum unstable in other. We elaborately discuss this feature below with some benchmark points.

We have selected two benchmark points (BP1 and BP2) as samples out of the allowed regions consistent with the relic density constraints. These points demonstrate how the different theoretical constraints can truncate the scale of validity of this scenario for two different values of right-handed neutrino masses M=110 TeV and $M=10^9$ TeV respectively. The parameter values are listed in Table. 1. BP1 and BP2 yield $\Omega_{\rm DM}h^2=0.1151$ and 0.1207 respectively, which is within with the allowed range of relic density. For M=110 TeV, BP1 ensures a stable vacuum till the Planck scale (Fig. 7a). It is

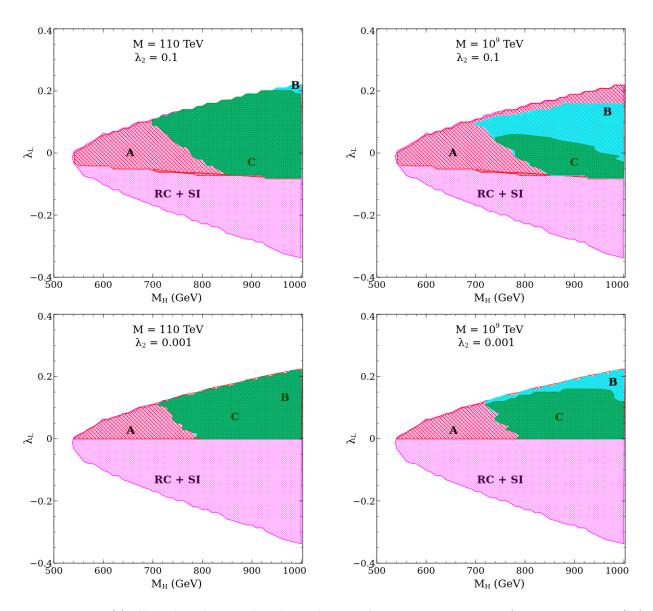


Figure 5: Region(s) allowed in the M_H - λ_L plane obeying the various constraints for M=110 TeV (left panel) and $M=10^9$ TeV (right panel) with three different choices of Λ_{UV} and two values of λ_2 (upper and lower panel). The full region (marked by $\mathbf{RC} + \mathbf{SI}$) (magenta) is allowed by the DM constraints alone. The overlapped regions labeled by \mathbf{A} (red), \mathbf{B} (cyan) and \mathbf{C} (green) are consistent with the theoretical constraints up to $\Lambda_{UV}=10^6$, 10^{16} and 10^{19} GeV, respectively.

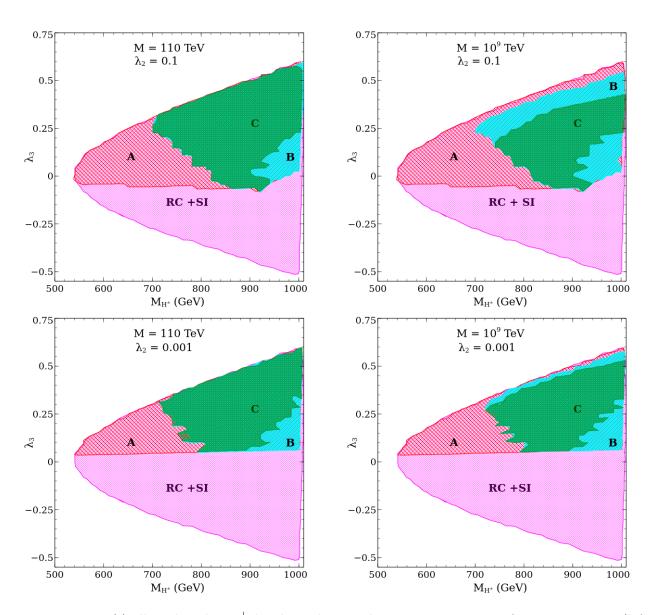


Figure 6: Region(s) allowed in the M_H^{\pm} - λ_3 plane obeying the various constraints for M=110 TeV (left panel) and $M=10^9$ TeV (right panel) with three different choices of Λ_{UV} and two values of λ_2 (upper and lower panel). The full region (marked by $\mathbf{RC}+\mathbf{SI}$) (magenta) is allowed by the DM constraints alone. The overlapped regions labeled by \mathbf{A} (red), \mathbf{B} (cyan) and \mathbf{C} (green) are consistent with the theoretical constraints up to $\Lambda_{UV}=10^6$, 10^{16} and 10^{19} GeV, respectively.

BP	M_H	M_{H^\pm}	M_A	λ_L	λ_2
BP1	850.0 GeV	$854.0~{\rm GeV}$	$858.0~{\rm GeV}$	0.02	0.1
BP2	710.0 GeV	712.0 GeV	711.0 GeV	0.11	0.1

Table 1: Benchmark values (BP) of parameters affecting the RG evolution of the quartic couplings. For each BP, two values of M, namely, 110 TeV and 10^9 TeV, have been used.

also consistent throughout with perturbativity and unitarity. However, one has $y_{\nu} = 0.168$ at $M = 10^9$ TeV. For this value, the term $\mathcal{O}(\lambda_2 y_{\nu}^2)$ has a dominant role in the RG evolution and λ_2 starts increasing rapidly from 10^9 TeV on wards (Fig. 7b).

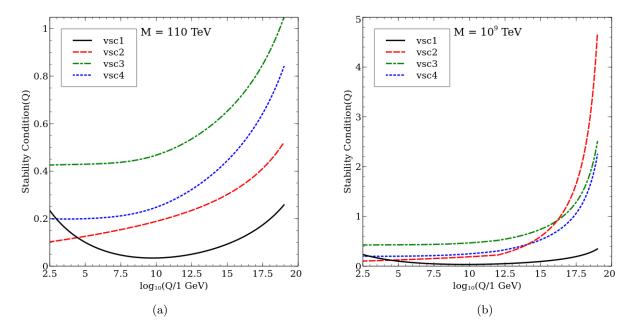


Figure 7: RG running of different scalar quartic couplings corresponding to BP1. The solid, dashed, dashed dotted and dotted lines denote the evolution curves of the stability conditions vsc1, vsc2, vsc3 and vsc4 respectively.

For M=100 TeV, BP2 exhibits similar features in the evolution trajectory as in BP1 (Fig. 8a). For $M=10^9$ TeV, the Yukawa coupling y_{ν} starts with an initial value around 0.51. This is accounted for by the very small mass splitting, of the order of a GeV, between H and A. Thus, the dominant contribution from the RH neutrinos comes from the $\mathcal{O}(y_{\nu}^4)$ term that causes λ_2 to drop sharply (Fig. 8b). Hence, in BP2, vacuum stability is destroyed shortly after 10^9 TeV as the condition $\lambda_2 > 0$ gets violated. This particular feature can only be witnessed in the case of closely spaced M_H and M_A , which is the primary requirement to satisfy the relic density constraints discussed before. This completes the explanation of how the allowed area can shrink due to different theoretical constraints for $M=10^9$ TeV.

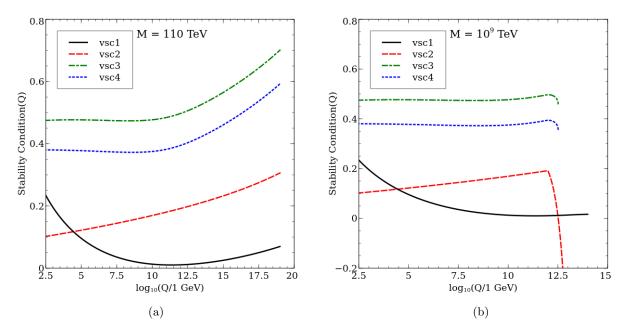


Figure 8: Same as Fig. 7 but for the benchmark point BP2.

7 Summary and Conclusions

We have examined the high-scale validity of a scenario that (a) offers a scalar dark matter, (b) radiatively generates Majorana masses for neutrinos, and (c) is responsible for leptogenesis. For this, we extend the SM fields with one additional inert Higgs doublet field (Φ_2) and three right handed neutrinos (N_i). These new particles are odd under a discrete Z_2 symmetry, while all the SM particles are even. Because of this discrete symmetry, Φ_2 does not acquire any vacuum expectation value (vev) and has no tree-level couplings to fermions. In this scenario, one has five physical scalars (h, H, A, H^{\pm}), where, h is denoted as the SM like Higgs boson with a mass of 125 GeV. The lightest state between H and A is the dark matter candidate due to built in Z_2 symmetry. In our analysis we have assumed H to be the dark matter candidate. Neutrino masses are generated at the one-loop level. The neutrino masses and mixing angles are determined in terms of Yukawa couplings (y_{ν}), new Higgs particle masses (M_H, M_A) and three heavy Majorana masses ($M_{1,2,3}$). In our numerical analysis we have assumed M_1 is mass of the lightest state and considered two values, namely, $M_1 \equiv M = 110$ TeV and 10^9 TeV. These two mass scales are consistent with Leptogenesis . For simplicity, in our analysis, we have considered only one diagonal Yukawa coupling and to determine the value of this coupling we have scanned over M_H and M_A for a given value of M, by keeping $M_{\nu} \sim \mathcal{O}(0.1 \text{ eV})$.

The parameter space of this model is determined in terms of additional Higgs boson masses, M_H , M_A , $M_{H^{\pm}}$, one quartic coupling λ_2 and a set of quartic coupling combinations, λ_L . While scanning the parameter space of this model, we have imposed the LEP bound on scalar masses, M_H , M_A and $M_{H^{\pm}}$.

As far as the dark matter constraints are concerned, we have used the relic density limits obtained at 3σ uncertainty from the PLANCK experiment and the direct detection cross-section limit from the

LUX experiment. Finally, for $M_H < M_h/2$, which would lead to large invisible decay width of the SM like Higgs boson, we demanded that the corresponding branching ratio is less than 20% as obtained from the model independent Higgs precision analysis.

With these set of constraints in hand, we have then scanned the IDM parameter space for two different ranges of dark matter masses, 50 GeV $< M_H < 90$ GeV and $M_H > 500$ GeV. It should be noted that with M_h fixed at 125 GeV, λ_2 , λ_L , M_H , M_A and M_{H^\pm} determined all the remaining quartic couplings. We have used these quartic couplings as the electroweak boundary condition by setting the starting RG running scale $Q = M_t$ and evolved these couplings up to the scale $\Lambda_{\rm UV}$, where this scenario remained consistent. In the RG evolution of these quartic couplings, the neutrino Yukawa couplings started playing its role from the right handed neutrino mass scale Q = M on wards. Following are the salient features of this model that our analysis brings out.

- In our study we have explicitly demonstrated that at the low DM mass region, the vacuum stability, perturbativity and unitarity constraints put stringent limits on the low-energy value of the coupling λ_L and the Z_2 -odd scalar masses M_{H^\pm}, M_A . These bounds strongly depend upon the scale up to which the theory is valid and the right-handed neutrino mass scale. It is interesting to note that all the parameter space allowed by the DM relic density and direct detection constraints lies well within the region allowed by the theoretical constraints valid up to the Planck scale. However, once we have imposed the condition that the Higgs to diphoton signal strength $(\mu_{\gamma\gamma})$ should lie within 2σ of the weighted average value of the ATLAS and CMS data, the allowed parameter space further squeezed.
- The scenario with high DM mass region ($M_H > 500$ GeV) is significantly different from that of the low DM mass region. In this case, the DM being very heavy, the constraints from direct detection is rather insignificant. On the other hand, the relic density constraint is ensured by a degenerate Z_2 -odd scalars ($\Delta M \simeq 10$ GeV). As a result of these, a large part of the parameter space in $\lambda_L M_H$ and $\lambda_3 M_{H^{\pm}}$ planes remain unconstrained.
- However, the study of the high scale validity of high DM mass region has interesting consequences. The DM-allowed region reduces substantially after imposing the theoretical constraints and this reduction is strongly dependent on the cut-off scale ($\Lambda_{\rm UV}$). The effect of neutrino Yukawa couplings in the RG evolution of the different quartic couplings are also prominent in this case. We have found that with the increase in the right handed neutrino mass scale M, the neutrino Yukawa coupling (y_{ν}) also increases, which in turn further reduces the allowed parameter space by either destabilizing the vacuum or violating the perturbativity bound. There is nonetheless a clearly identifiable region of the parameter space, which keeps the model valid all the way up to the Planck scale. This scenario is consistent with the measured Higgs-to-diphoton rates as measured during the 8 TeV run of the LHC.

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Appendix

A One-loop Renormalization group (RG) equations

The RG equations for the gauge couplings, for this model, are given by [42],

$$16\pi^2 \frac{dg_s}{dt} = -7g_s^3, (23a)$$

$$16\pi^2 \frac{dg}{dt} = -3g^3, (23b)$$

$$16\pi^2 \frac{dg'}{dt} = 7g'^3. (23c)$$

Here g', g and g_s denote the U(1), SU(2)_L and SU(3)_c gauge couplings respectively.

The quartic couplings λ_i (i = 1, ..., 5) evolve according to,

$$16\pi^{2} \frac{d\lambda_{1}}{dt} = 12\lambda_{1}^{2} + 4\lambda_{3}^{2} + 4\lambda_{3}\lambda_{4} + 2\lambda_{4}^{2} + 2\lambda_{5}^{2} + \frac{3}{4}(3g^{4} + g'^{4} + 2g^{2}g'^{2}) -\lambda_{1}(9g^{2} + 3g'^{2} - 12y_{t}^{2} - 12y_{b}^{2} - 4y_{\tau}^{2}) - 12y_{t}^{4},$$
(24a)

$$16\pi^{2} \frac{d\lambda_{2}}{dt} = 12\lambda_{2}^{2} + 4\lambda_{3}^{2} + 4\lambda_{3}\lambda_{4} + 2\lambda_{4}^{2} + 2\lambda_{5}^{2} + \frac{3}{4}(3g^{4} + g'^{4} + 2g^{2}g'^{2}) - 3\lambda_{2}(3g^{2} + g'^{2} - \frac{4}{2}y_{\nu}^{2}) - 4y_{\nu}^{4},$$
(24b)

$$16\pi^{2} \frac{d\lambda_{3}}{dt} = (\lambda_{1} + \lambda_{2}) (6\lambda_{3} + 2\lambda_{4}) + 4\lambda_{3}^{2} + 2\lambda_{4}^{2} + 2\lambda_{5}^{2} + \frac{3}{4} (3g^{4} + g'^{4} - 2g^{2}g'^{2}) -\lambda_{3} (9g^{2} + 3g'^{2} - 6y_{t}^{2} - 6y_{b}^{2} - 2y_{\tau}^{2} - 2y_{\nu}^{2}),$$
(24c)

$$16\pi^{2} \frac{d\lambda_{4}}{dt} = 2(\lambda_{1} + \lambda_{2}) \lambda_{4} + 8\lambda_{3}\lambda_{4} + 4\lambda_{4}^{2} + 8\lambda_{5}^{2} + 3g^{2}g^{2}$$
$$-\lambda_{4}(9g^{2} + 3g^{2} - 6y_{t}^{2} - 6y_{b}^{2} - 2y_{\tau}^{2} - 2y_{\nu}^{2}), \qquad (24d)$$

$$16\pi^2 \frac{d\lambda_5}{dt} = (2\lambda_1 + 2\lambda_2 + 8\lambda_3 + 12\lambda_4)\lambda_5 - \lambda_5(9g^2 + 3g'^2 - 6y_b^2 - 2y_\tau^2 - 6y_t^2 - 2y_\nu^2), \quad (24e)$$

For the Yukawa couplings the corresponding set of RG equations are,

$$16\pi^2 \frac{dy_b}{dt} = y_b \left(-8g_s^2 - \frac{9}{4}g^2 - \frac{5}{12}g'^2 + \frac{9}{2}y_b^2 + y_\tau^2 + \frac{3}{2}y_t^2 \right), \qquad (25a)$$

$$16\pi^2 \frac{dy_t}{dt} = y_t \left(-8g_s^2 - \frac{9}{4}g^2 - \frac{17}{12}g'^2 + \frac{9}{2}y_t^2 + y_\tau^2 + \frac{3}{2}y_b^2 \right), \tag{25b}$$

$$16\pi^2 \frac{dy_{\tau}}{dt} = y_{\tau} \left(-\frac{9}{4}g^2 - \frac{15}{4}g'^2 + 3y_b^2 + 3y_t^2 + \frac{1}{2}y_{\nu}^2 + \frac{5}{2}y_{\tau}^2 \right). \tag{25c}$$

$$16\pi^2 \frac{dy_{\nu}}{dt} = y_{\tau} \left(-\frac{9}{4}g^2 - \frac{3}{4}g'^2 - \frac{3}{4}y_{\tau}^2 + \frac{5}{2}y_{\nu}^2 \right). \tag{25d}$$

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